

SW Ar, Kr, and Xe from Bulk Collector Array CZ-Si targets

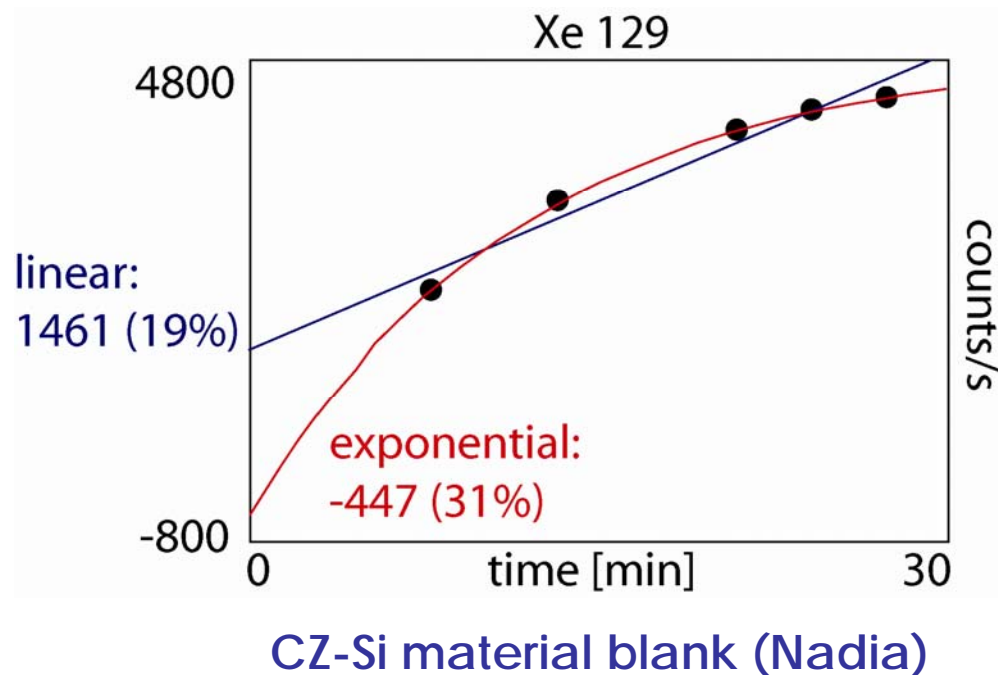
Talk: Wednesday 1:45 pm
Vogel N., Heber V.S., Baur H., Burnett D.S., Wieler R.

Outline

- State of the project Aug/Sept 2008; memory problem
- Data
- Comparison of our data set to:
 - other Genesis AKX data,
 - lunar SW record,
 - solar photosphere data

State of project Aug/Sept 08

- 5 successful UV-laser ablation AKX analyses of CZ-Si targets by V. Heber
- Spectrometer memory problem:



Non-linear increase of
gas amounts (KX) →
regression impossible!



Unknown species
causes memory

⇒ first task: specification & elimination of unknown species

What causes SMP?

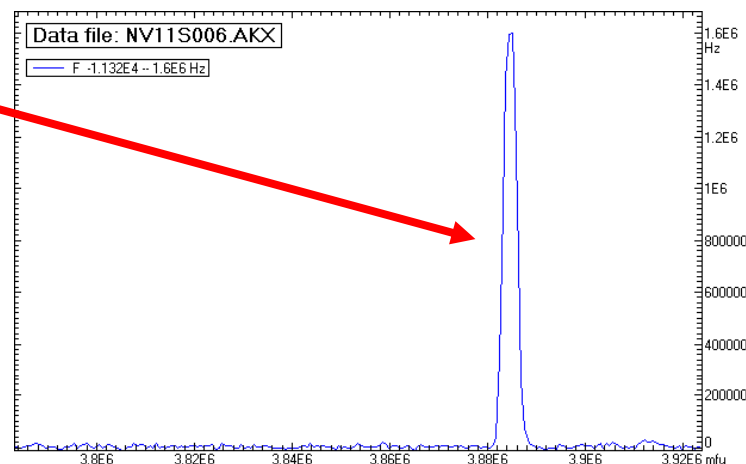
Search for culprit (Nov. 2008):

- Unusual peaks on masses **19 (F)**, 34.5, **69**

⇒ **Memory causing ion: $(\text{CF}_3)^+$ (?)**

Potential source: teflon $(\text{C}_2\text{F}_4)_n$?

- not removed by heating
- cracked by UV-laser light



Source elimination:

- Unsuccessful: memory worse with each test (different blank material, viewport cleaning, viewport, sample holder, preheating....)

Different gas cleaning:

- Unsuccessful: mass 69 behaves identical to Kr

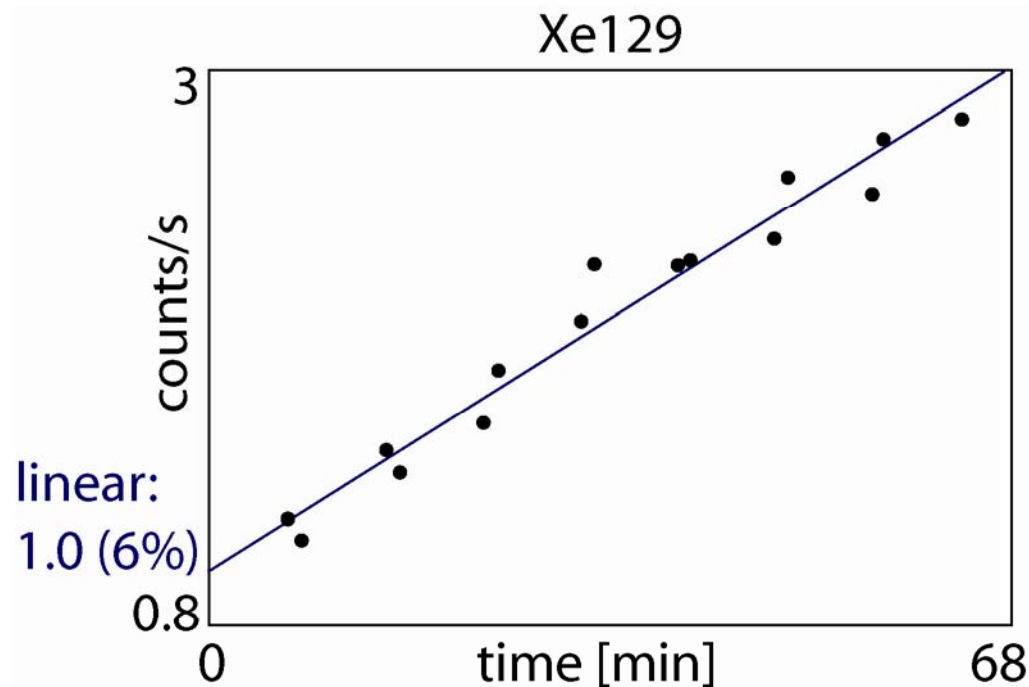
Solution of SMP (christmas 08)

⇒ contaminant from oil in pressurized air despite filter?

⇒ Yes; moreover: filter made of teflon, old, crumbly

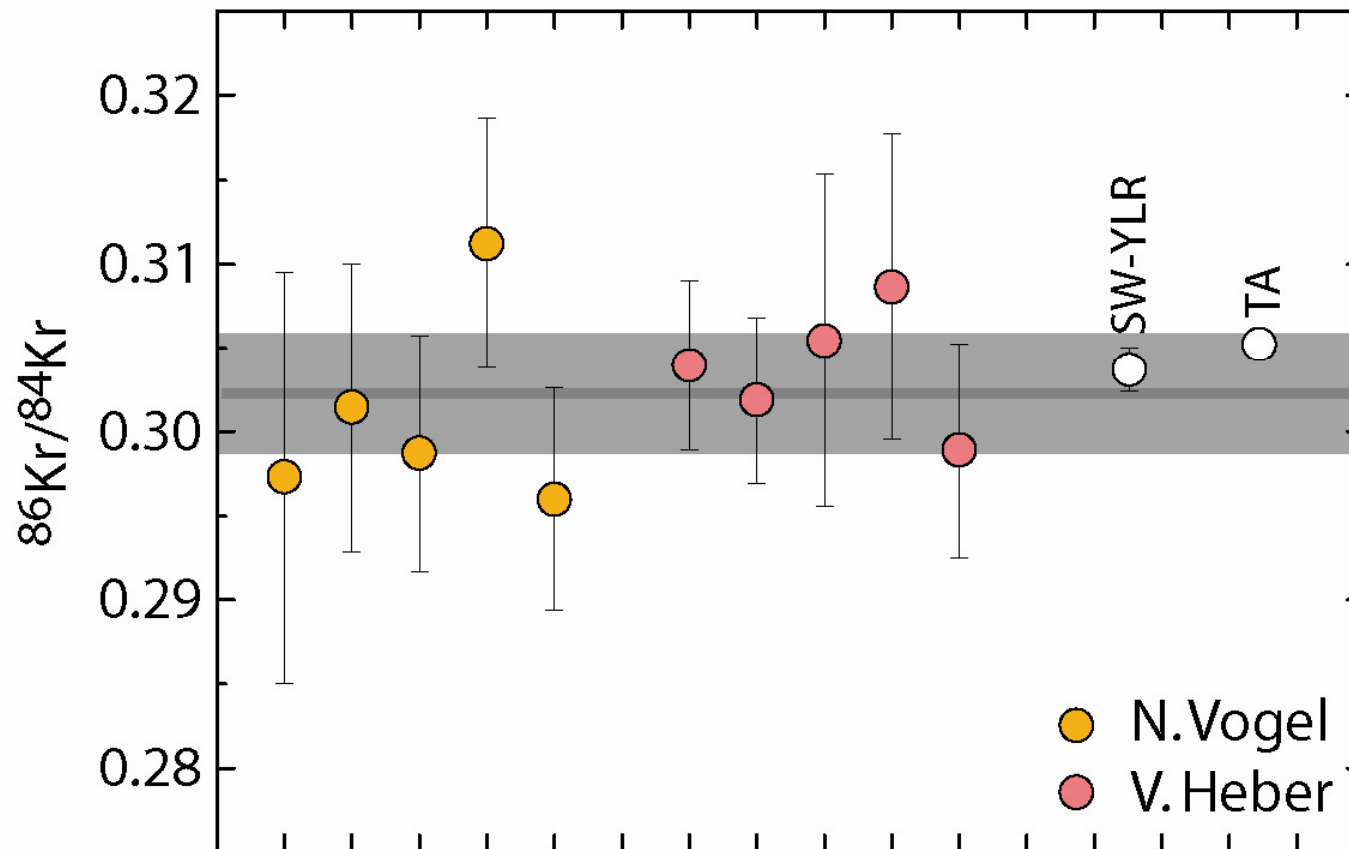
Procedure:

- no use of pressurized lab air for cleaning,
- exchange of contaminated chamber parts



CZ-Si material blank (10 mm²)

Finally - data! Kr isotopic composition



- Vogel & Heber identical
- **Average (n=10):**
 0.302 ± 0.004
- Identical to lunar SW record

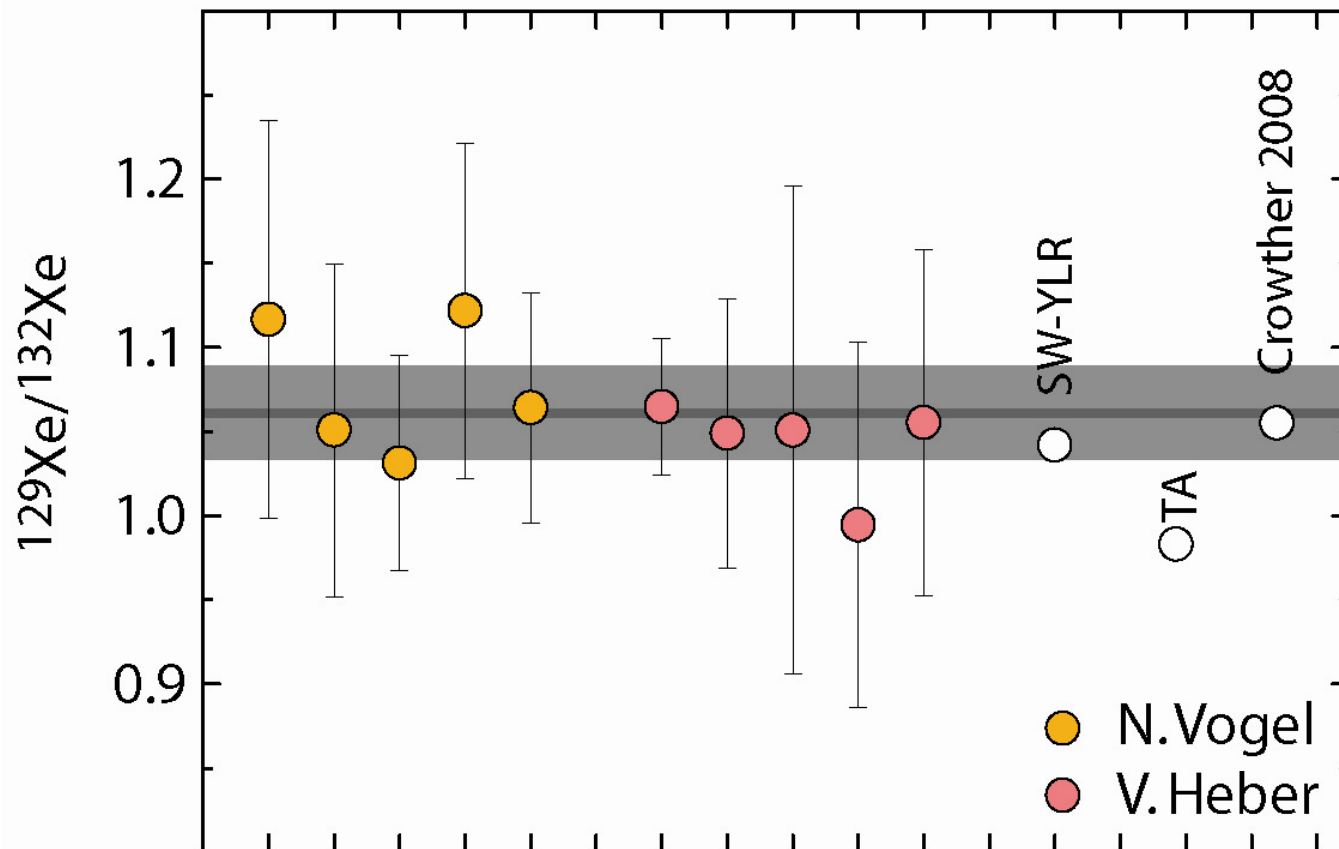
YLR = **Y**oung **L**unar **R**egolith

TA = **T**errestrial **A**tmosphere

SP= **S**olar **P**hotosphere

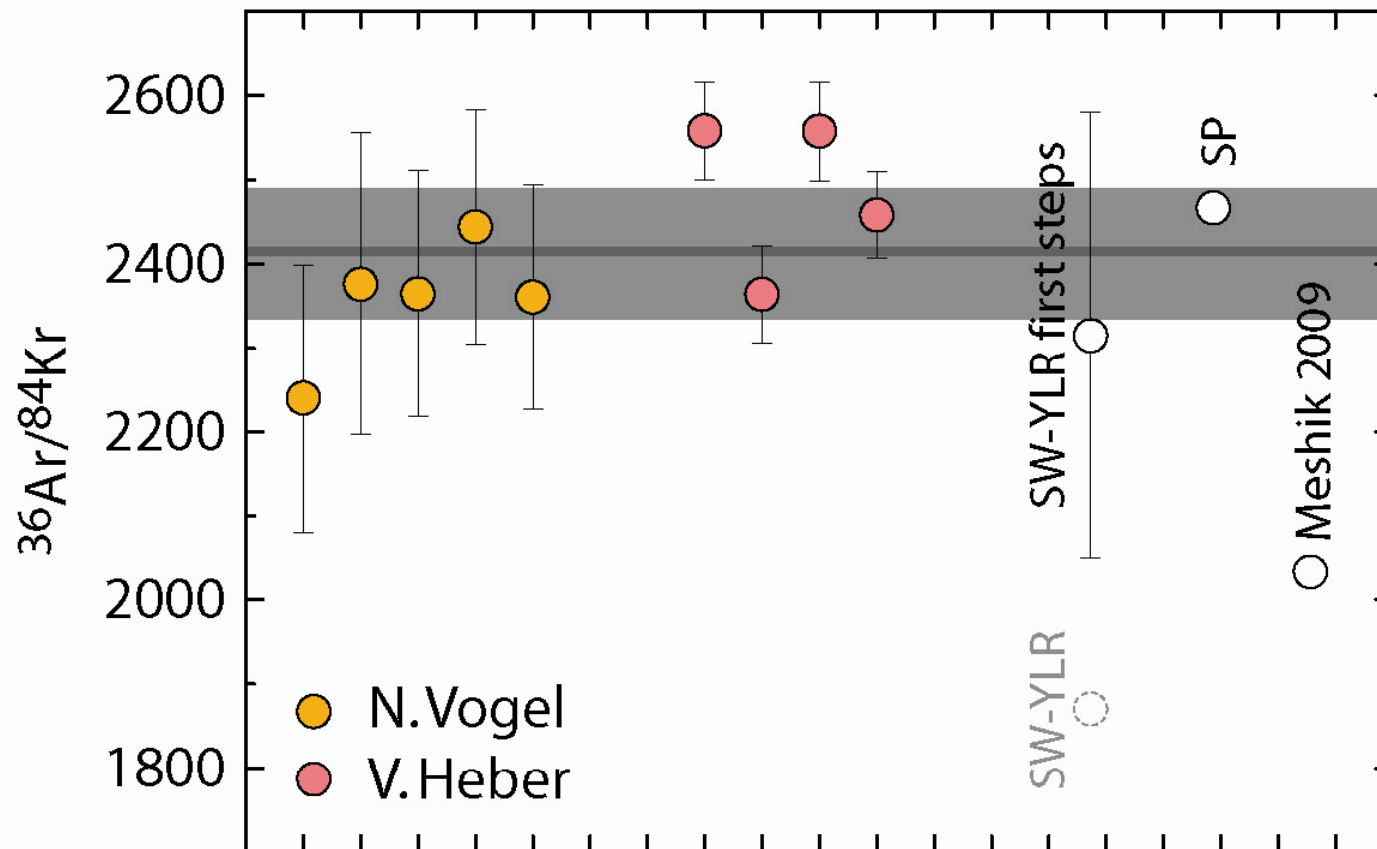
Averages with standard errors of the mean (95% confidence level)

Xe isotopic composition



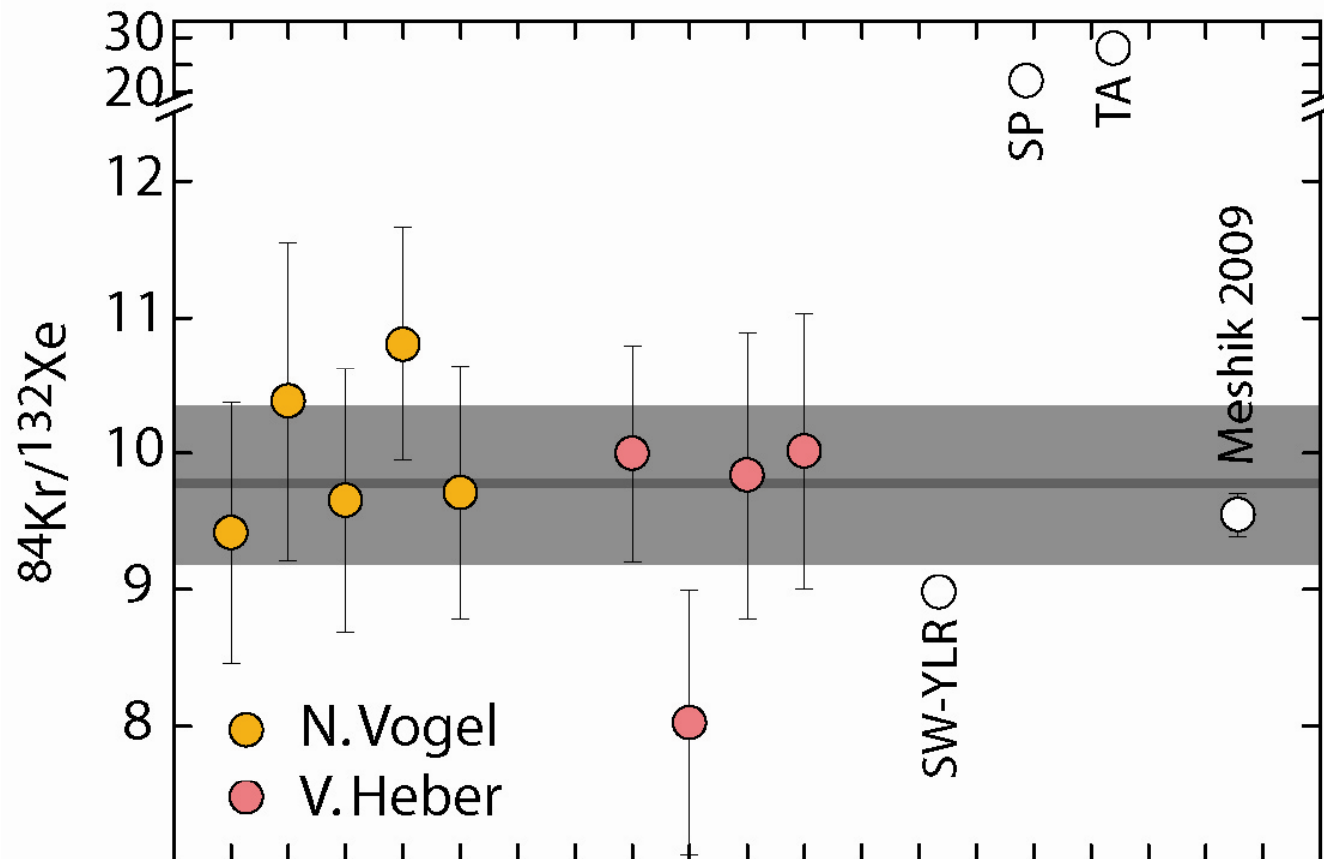
- Vogel & Heber identical
- **Average (n=10):**
 1.06 ± 0.03
- Identical to lunar SW record
- Identical to Genesis data by Crowther et al. 2008 (RELAX)

Ar/Kr elemental ratio



- Difference Heber & Vogel → blank
- **Average (n=9): 2414 ± 77**
- Agreement with SW lunar record?
- Agrees with SP (Lodders 03/08)
- Low Meshik et al. 2009 Ar/Kr „Different Ar, Kr implantation“

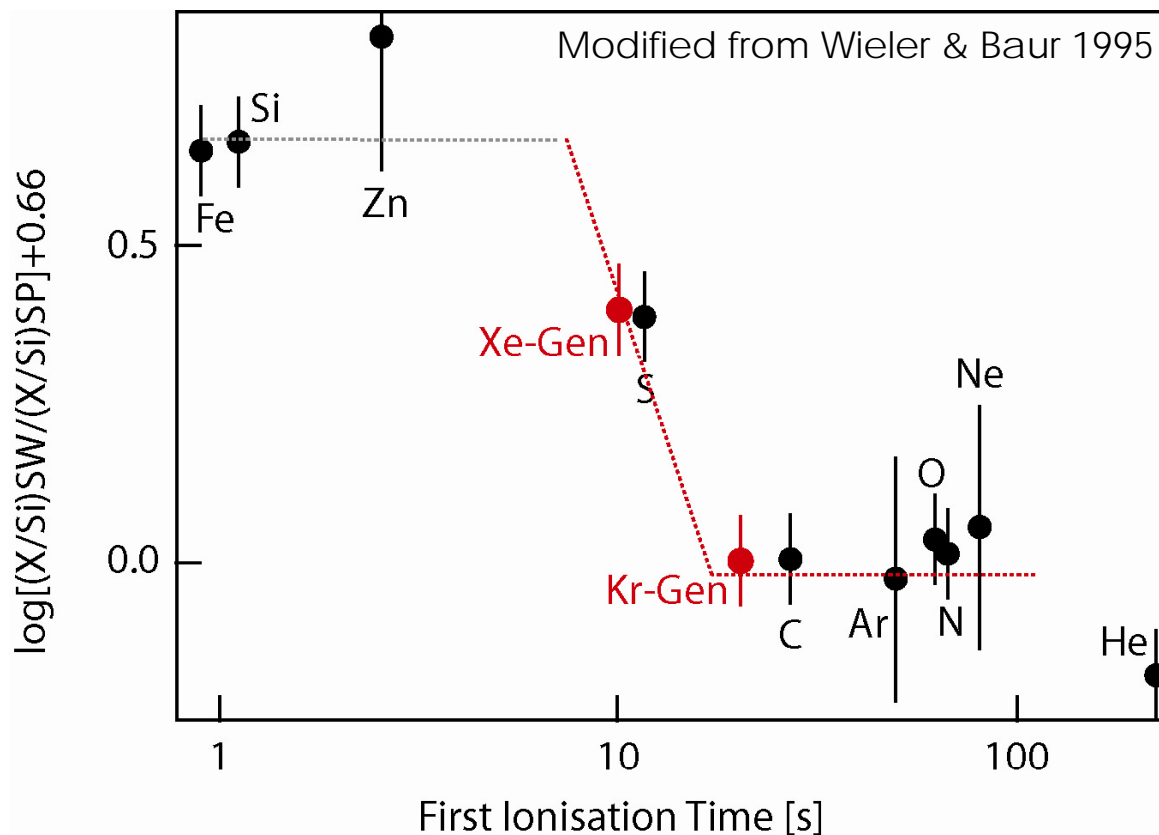
Kr/Xe elemental ratio



- Vogel & Heber identical
- **Average (n=9):**
 9.8 ± 0.6
- Identical to lunar SW record & Meshik 2009
- Clearly different from SP → **element fractionation**
Sun - SW

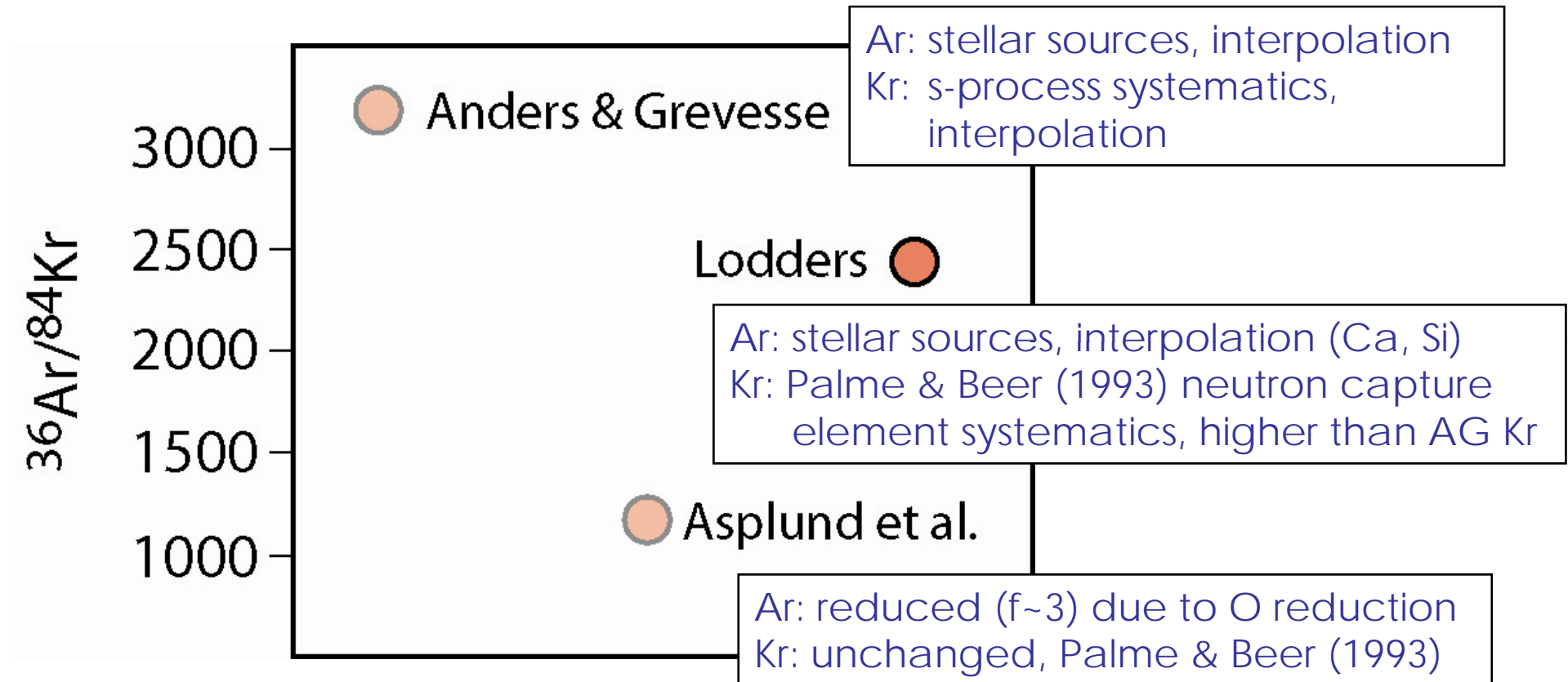
Element fractionation Sun - SW

Element fractionation during ion-neutral separation in upper chromosphere according to FIP and FIT (low FIP/FIT elements enriched in SW (e.g., Geiss 1989))



- W & B (1995):
Kr: partial enrichment (x1.8) in SW
Xe: enrichment = low FIT elements (x4) in SW
- SW Genesis, new SP data:
Kr: no enrichment
Xe: partial enrichment (x2.2)
- dependence on SP values, which are still not settled (e.g., Ar/Kr)

Example: solar photospheric Ar/Kr



- ⇒ Lodders 2003/8: most reliable SP Ar/Kr
- ⇒ agreement with new Genesis data strengthens both ratios

Conclusions

- Don't use pressurized air...
- Genesis provides ~consistent data base for modern SW Ar, Kr, Xe
(Zürich, Manchester, St. Louis)

Isotopic, elemental ratios:

$^{86}\text{Kr}/^{84}\text{Kr}$:	0.302 ± 0.004
$^{129}\text{Xe}/^{132}\text{Xe}$:	1.06 ± 0.03
$^{36}\text{Ar}/^{84}\text{Kr}$:	2414 ± 77
$^{84}\text{Kr}/^{132}\text{Xe}$:	9.8 ± 0.6

Ar, Kr, Xe fluences [atoms/cm²]:

^{36}Ar :	$(2.95 \pm 0.05) \times 10^{10}$ [n=10]
^{84}Kr :	$(1.22 \pm 0.04) \times 10^7$ [n=9]
^{132}Xe :	$(1.26 \pm 0.11) \times 10^6$ [n=9]

- Agreement with SW record from young lunar regoliths
- Compared to SP no enrichment of Kr, partial enrichment of Xe in SW
(SP values might not be as settled as they appear!)
- Final measurements: 2-3 more samples + several large material blanks
(significant source of uncertainty)



Origin of solar photospheric values

KX: based on theoretical values from neutron-capture systematics

Kr (in Lodders 2003) from Palme & Beer (1993) who consider contributions from main and weak s-process to ^{82}Kr + isotopic composition of Wieler 2002 to derive Kr elemental abundance (13% higher than the value of A & G, in line with Raiteri et al (1993)) "A&G Kr value underestimated by ~20%

Xe (in Lodders 2003) most recent measured neutron cross sections for important Xe process nuclei (Reifarth et al. 2002) + Wieler et al 2002 isotopic ratios; elemental abundances 10-20 % higher than previously, much smaller uncertainties.

→ Different sources come up with K/X ratios of max 20% difference: does not explain factor of >2 for difference between SW – SP ratio